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Workshop 6-13 June 1998, Bern, Switzerland (Space
Sciences Series Of ISSI)**

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Space Sciences Series of ISSI Volume 7, 1999, pp 21-52 Corotating interaction regions are a consequence of spatial variability in the coronal .. Regions; Book Subtitle: Proceedings of an ISSI Workshop 6-13 June 1998, Bern, Switzerland

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May 30, 2007 A.B. in Physics awarded in June 1965. 2003 Thomas M. Donahue Collegiate Professor of Space Science . 6/1979 - 4/1987 NASA, Solar Wind Ion Composition Spectrometer on .. pick-up ions in corotating interaction regions, Geophys. Proc. of an ISSI Wksp., Bern, Switzerland, 26-30 January, 1998,.

Nasa-cr-202513

constant. Corotating interaction regions were first identified as recurrent increases in H and He fluxes with 27-day period in 1965 solar minimum in

Cir - wikipedia, the free encyclopedia

Corotating Interaction Region, the interface between fast and slow regions of the solar wind; Cosmetic Ingredient Review, a consumer safety group;

Hao 2011 profiles in science: dr. alan burns

Corotating Interaction Region / High Speed Stream Storm Effects on the Thermosphere and Ionosphere Burns, A.G., S.C. Solomon, L. Qian, W. Wang,

Elemental composition of the inner source pickup

The acceleration of charged particles in corotating interaction regions Proceedings of the International Astronomical Union, 2008, 4, S257, 17CrossRef;

Corotating interaction regions - proceedings of

A Corotating Interaction Region (CIR) is the result of the interaction of fast solar wind with slower solar wind ahead. CIRs have a very large

Corotating interaction regions - harvard

Abstract Corotating Interaction Regions are long lasting large-scale plasma structures generated in low and middle latitude regions of the heliosphere by the

Compression of jupiter s magnetosphere by the

Compression of Jupiter s magnetosphere by the solar wind Reexamination via MHD simulation of evolving corotating interactions regions (1981)

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245 10 Corotating interaction regions : \$b proceedings of an ISSI. Workshop, 6- 13 June 1998, Bern, Switzerland / \$c ed. by A. Balogh [et al.] 260 Dordrecht 830 0 Space science series of ISSI \$v 7 \$w 1120178. ERKL RUNGEN. Nur wenn

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Chamberlin, P., CCMC results shown on MAVEN Project Science Group as an example of minimum, Invited Oral Paper, ISSI Workshop, Long-term Reconstruction of Solar and Solar Wind Parameters, Bern, Switzerland, May 2012. .. K., Galvin, A.B., Zhang, T.L., Initiation and evolution of corotating interaction regions,Eos

Introduction: the existence of the heliosphere and its main

The heliosphere is the vast region of space surrounding the Jokipii, Simnett, von Steiger and Wenzel (1998); and on corotating interaction regions, edited by

13 variability of low energy cosmic rays near

Mar 21, 2012 1998; Cane et al., 1999 among others) and it is examined further along with diffusion barriers were identified as merged interaction regions, MIRs . Page 6 Proceedings of an ISSI Workshop 21-26 March 1999, Bern,. Switzerland, Series: Space Sciences Series of ISSI, Vol. event on June 3, 1982.

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Chief editor for SOHO 23 workshop proceedings (ASP Conf. . Cranmer, S. R. 2009, Coronal Holes, Living Reviews in Solar Physics, 6, 3. . Cranmer, S. R., and Owocki, S. P. 1996, Hydrodynamical Simulations of Corotating Interaction Regions . Science Institute (ISSI), January 25 29, 2010, Bern, Switzerland (Invited

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Corotating interaction regions -

Corotating Interaction Regions are long lasting large-scale plasma structures generated in low and middle latitude regions of the heliosphere by the interaction

Forecasting the dst index during corotating

Forecasting the Dst index during corotating interaction region events using synthesized solar wind parameters. Authors. T. Andriyas, E-mail address:

Corotating interaction regions during the recent

The declining phase of solar activity cycle 23 has provided an unprecedented opportunity to study the evolution and properties of corotating interaction regions

The three-dimensional morphology of a corotating

the three-dimensional morphology of a corotating interaction region in the inner heliosphere free issue. b. e. wood, r. a. howard, a. thernisien, and d. g

Solar phenomena - wikipedia, the free encyclopedia

Space weather disturbances can cause solar storms on Earth, disrupting . A video of the series of coronal mass ejections in August 2010 .. "Dr Karl's Great Moments In Science: Lazy Sun is less energetic than compost". . Regions, Corotating Interaction Regions Proceedings of an ISSI Workshop, 6 13 June 1998, Bern,

Abstract - harvard university

Abstract We have examined the relationships among coronal holes (CHs), corotating interaction regions (CIRs), and geomagnetic storms in the period 1996 - 2003.

Comparison of the heliospheric magnetic field in

from 20 to 40 S, a region previously explored at a time of declining solar activity Corotating Interactions are the predominant feature in both data sets but are

Corotating interaction regions (space sciences

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Geomagnetic storm - wikipedia, the free

Several space weather phenomena tend to be associated with or are caused A geomagnetic storm is defined by changes in the Dst (disturbance storm time) index. .. Corotating Interaction Regions, Corotating Interaction Regions Proceedings of an ISSI Workshop, 6 13 June 1998, Bern, Switzerland, Springer (2000),

Numerical heliospheric simulations as assisting

Numerical Heliospheric Simulations as Assisting Tool for Interpretation of Observations We illustrate that corotating interaction regions (CIRs) show up differently

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and the co-rotating interaction regions they engender, the interaction with increased solar wind speed amplifies the magnetosphere-ionosphere-thermosphere

Geoeffectiveness of corotating interaction regions

Geoeffectiveness of corotating interaction regions as measured by Dst index M. V. Alves,¹ E. Echer,² and W. D. Gonzalez² Received 23 August 2005; revised 23 October

Ptys/lpl faculty | lunar and planetary laboratory

Major current thrusts revolve around work on the Voyager and ACE space missions, . Proceedings. ISSI Workshop, Bern (Switzerland), 6 - 13 Jun 1998. ISSI (International Space Science Institute) Workshop, Bern (Switzerland), 17 - 20 Sep Kota, J., Jokipii, J. R. The role of corotating interaction regions in cosmic- ray

Potchefstroomse universiteit vir c - north-west

13. Hattingh, M., & Burger, R.A., Potgieter, M.S., & Haasbroek, L.J. 1997. Modulation of Galactic Cosmic Rays at Solar Minimum, Space Science Corotating interaction regions in the outer heliosphere - Report of Working Group 4, .. June 2012. . 6. May 1999. Discussions with Dr. H. Fichtner. ISSI, Bern, Switzerland. 7.

Multi-point observations of corotating interaction

PROCEEDINGS OF THE 31 ICRC, OD Z 2009 1 Multi-Point Observations of Corotating Interaction Regions from STEREO and ACE R. A. Leske , G. M. Mason

First imaging of corotating interaction regions

First imaging of corotating interaction regions using the STEREO spacecraft A. P. Rouillard,^{1,2} J. A. Davies,² R. J. Forsyth,³ A. Rees,³ C. J. Davis,² R. A. Harrison,²

Media responsible for faraday rotation: a review -

[5] The science objectives of many of the new generation LF arrays, however, . of a point in a three dimensional space; s is a vector coordinate measured in m The coronal structures of most interest are the Corotating Interaction Regions ISSI Workshop, 6 13 June 1998, Bern, Switzerland, Kluwer Acad., Dordrecht,

Voyager 2 observations of corotating interaction

Space Sciences Series of ISSI Volume 7, 1999, pp 53-59. Voyager 2 Observations of Corotating Interaction Regions (CIRs) in the Outer Heliosphere.

Far ultraviolet spectroscopic explorer snapshot

Rybicki 1988); (2) the interfaces of large-scale corotating interactions regions (CIRs), high-energy processes in the optical/UV region of the spec-trum.

Corotating interaction region

CIR, Corotating Interaction Region, corotating interaction regions, corotating interacting regions, corotation interaction region, corotational interaction.

On the evolution of coronal mass ejections in the

Complexities in the interaction of the interplanetary medium and the ICME are generally ignored due to the or the passage of corotating interaction regions

Formation and evolution of corotating interaction

Corotating interaction regions are a consequence of spatial variability in the coronal expansion and solar rotation, which cause solar wind flows of different

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Propagation of corotating interaction region

Propagation of corotating interaction region. Date: 07 March 2012 Satellite: A CIR is a compression region between fast- and slow-moving solar wind streams.

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